



**Issues with convection: What is a useful framework beyond bulk models  
of large  $N$ , non-interacting, scale-separated, equilibrium systems?**

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# Outline



- Introduction: current treatments of convection
- Concerning bulk models
- Concerning  $N \nearrow \infty$
- Concerning spatial correlations
- Concerning time-dependence and equilibrium
- Summary



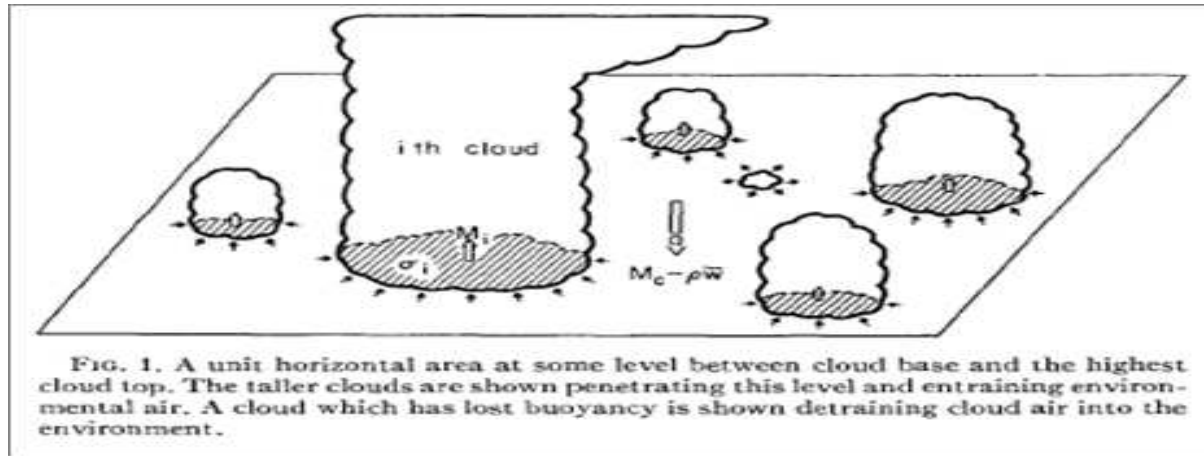


# Current treatments of convection



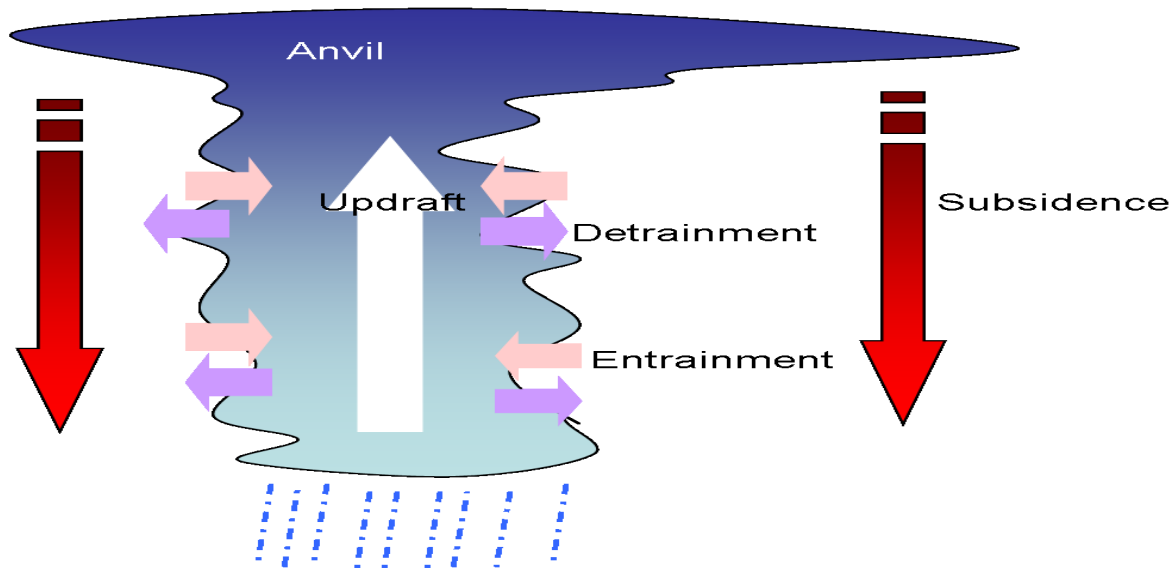
# The cumulus ensemble

The Arakawa and Schubert (1974) picture



- Convection characterised by ensemble of cumulus clouds
- Scale separation in both space and time between cloud-scale and the large-scale environment

# Entraining/detraining “plume”



Key variable is the mass flux,

$$M_i = \rho \sigma_i w_i$$

$$\overline{\rho \chi' w'} \approx \sum_i M_i (\chi_i - \chi_{\text{env}})$$

# Equations for a plume



- Various plume models based on this picture
- Differ in formulation of entrainment/detrainment and microphysics
- Integrate from cloud base up to terminating level where the in-cloud buoyancy vanishes
- Hardest part is the lower-boundary condition, or closure: i.e., what is the mass flux at cloud base for each  $i$ ?



# Closure



- The convection is being forced by some large-scale processes that act to destabilize the atmosphere
- If convection occurs, it will act to try to restore stability
- At equilibrium, the large-scale and convective tendencies are in balance



# Concerning bulk models





# Bulk parameterizations



- A more common approach in practice (MetUM, ECMWF, WRF...)
- Start from the plume equations, and sum over plumes
- Get back essentially the same equations with in-plume values replaced by bulk values,

$$\chi_B = \frac{\sum_i M_i \chi_i}{\sum_i M_i}$$

Just one “bulk plume” now, so all is much simpler...



# The price of a bulk scheme

A bulk method works because the plume equations are (almost!) linear

- Hinges on an extra “gross assumption” about the detraining cloud liquid water  
assumed equal to bulk value, which means condensate detrainment is systematically overestimated in a bulk model
- Linearity is needed in the microphysics and radiation terms  
By construction, cumulus microphysics and cumulus-radiation interactions are supposed to be very crude
- No simplification occurs for chemical transports



Concerning  $N \not\rightarrow \infty$



# Typical $N$ values



- Convective instability is released in discrete events
- A typical mass flux for one cloud is  $\sim 10^7 \text{ kgs}^{-1}$
- To stabilize a typical convective forcing in the tropics needs a total mass flux of  $\sim 10^{-2} \text{ kgm}^{-2} \text{ s}^{-1}$
- So a typical number of clouds is  $\sim 10^{-9} \times \text{area}$
- $\sim 10$  for a typical “grid box” of area  $(100\text{km})^2$

⇒ The number of clouds in a GCM grid-box is not large enough to produce a steady response to a steady forcing

e.g. Xu et al 1992; Shutts and Palmer 2004



# Variability



- Let's retain the equilibrium assumption, which determines total cloud-base mass flux required on average,  $\langle M \rangle$
- Want to describe variability arising from fluctuations about equilibrium
- Must consider the partitioning of  $\langle M \rangle$  into individual clouds
- i.e., we will need the pdf for the mass flux  $m$  of a single cloud
- and the pdf for the number of clouds present



# pdf for $m$



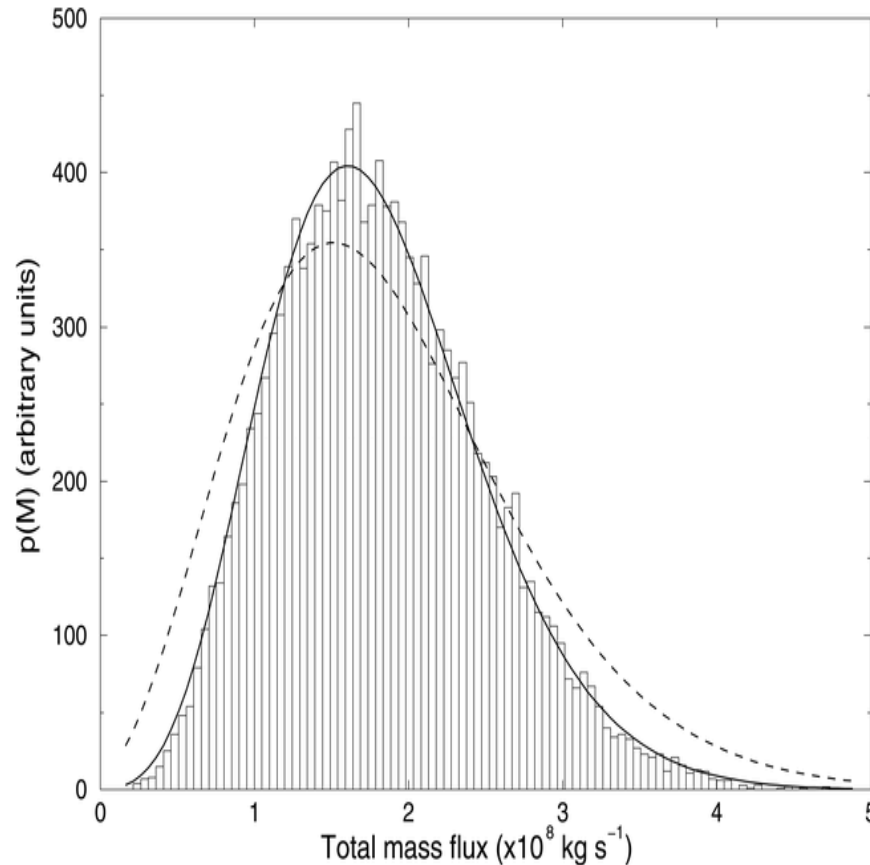
- Our assumptions about clouds as discrete, independent objects in a statistical equilibrium with a large-scale, macroscopic state are directly equivalent to those for an ideal gas
- So the pdf of  $m$  is a Boltzmann distribution

$$p(m)dm = \frac{1}{\langle m \rangle} \exp\left(\frac{-m}{\langle m \rangle}\right) dm$$

- Remarkably good and robust in CRM data  
Cohen and Craig 2006; Shutts and Palmer 2007; Plant and Craig 2008;  
Davies 2008; Davoudi et al 2010



# pdf for $M$



- Number of clouds is not fixed, unlike number of gas particles
- If they are randomly distributed in space, number in a finite region given by Poisson distribution
- pdf of the total mass flux is a convolution of this with the Boltzmann

# Stochastic parameterization



- Grid-box state  $\neq$  large-scale state  
space average over  $\Delta x \neq$  ensemble average
- We must parameterize convection on the grid-scale as being unpredictable, but randomly sampled from a known pdf dictated by the large-scale



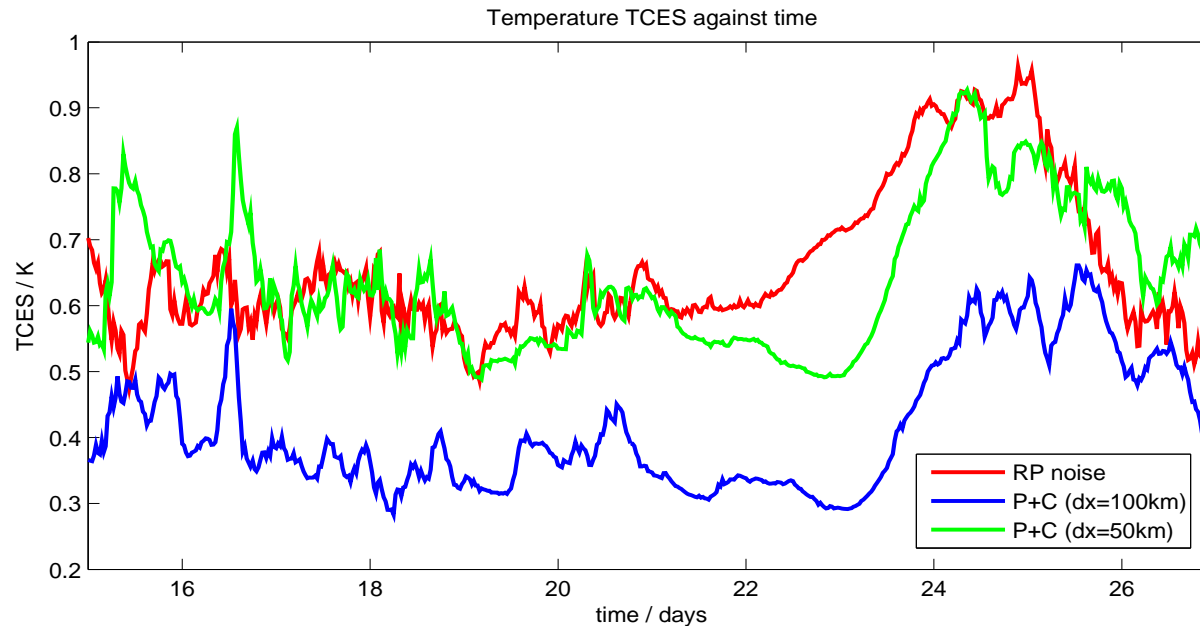
**Important note:** **None** of these scales is fixed in a simulation!





# Practical implementation

## Single-column test with Plant-Craig (2008) parameterization



- Spread similar to random parameters or multiplicative noise for  $\Delta x = 50\text{km}$
- Stochastic drift similar to changing between deterministic parameterizations



# Concerning spatial correlations



# Spatial correlation

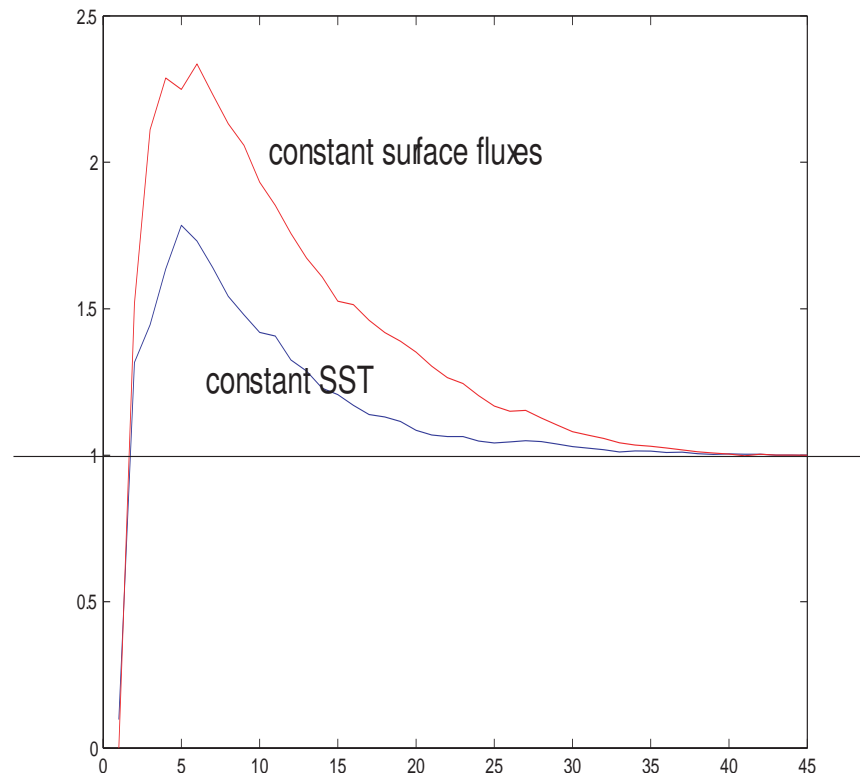


- We have assumed the clouds to be independent and randomly distributed in space
- In reality, they can readily (self-)organize, even in a uniform environment with uniform forcing
- This will affect both the mean response and the variability, but can we account for it?



# Non-random distribution

Consider  $w_{12}(r)$ , the expectation of finding a 2nd cloud a distance  $r$  from the 1st, normalized by that for a random spatial distribution



- Plot is in the equilibrium state of a CRM subject a forcing constant in space and time
- Self-organization with clumping at  $\sim 10$  to  $20\text{km}$

# Non-ideal gas analogy



- Statistical mechanics of gas particles easily generalized to include weak interactions between them
- First order correction is to consider only pairwise interactions between particles
- Each cloud is subject to an effective interaction potential

$$V_{12}(r) = -\langle m \rangle \ln w_{12}(r)$$



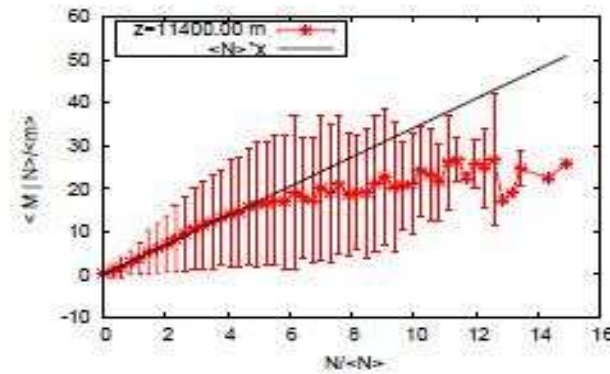
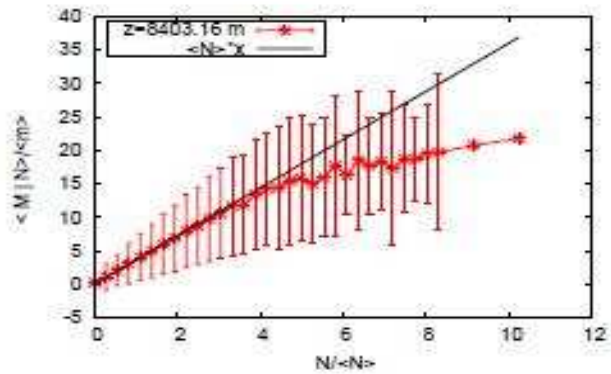
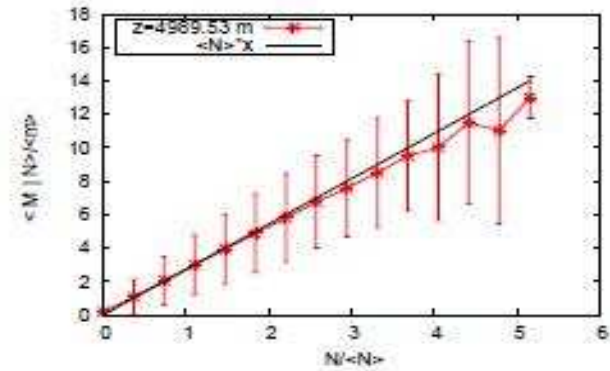
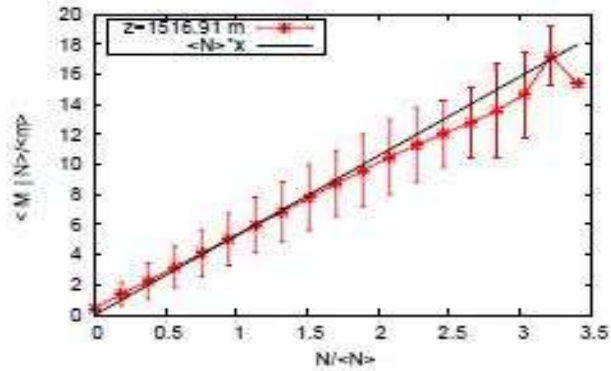
# What to expect?

$$\langle M|N \rangle = N \langle m \rangle \left[ 1 - \frac{N}{2A} \int 2\pi r (w_{12} - 1) dr \dots \right]$$

- Integral is  $> 0$  for clumping
- Deviations will be largest at large  $N/A$

# A hint?

CRM data from Davoudi 2008



$$\langle M | N \rangle = \langle m \rangle N$$

# Variance

$$\frac{\langle (\Delta N)^2 \rangle}{\langle N \rangle^2} = \frac{1}{\langle N \rangle} \left[ 1 + \frac{\langle N \rangle}{A} \int 2\pi r (w_{12} - 1) dr + \dots \right]$$

- Enhanced number variance if clouds clump together
- Plant-Craig parameterization makes a moderate underestimate of the variability
- **Could** parameterize organization straightforwardly from CRM experiments designed to study  $w_{12}$





# Concerning time-dependence and equilibrium



# Why consider time dependence?



- For relatively rapid forcings, we may wish to consider a prognostic equation for cloud-base mass flux  
Pan and Randall 1998; Piriou et al 2008
- Even for steady forcing, it is not obvious
  - that an equilibrium **must** be reached
  - which equilibrium might be reached



# Time dependence



- Let  $B_i$  be a vertical integral of in-cloud buoyancy over the depth of cloud  $i$
- After some algebra

$$\frac{dB_i}{dt} = F_i - \sum_j \gamma_{ij} M_j$$

where  $B$ ,  $F$  and  $\gamma$  are all calculable given a cloud model

- Also, the convective kinetic energy equation is

$$\frac{dK_i}{dt} = B_i M_i - \frac{K_i}{\tau_D}$$



# Closing this system



- Pan and Randall (1998) and others postulate

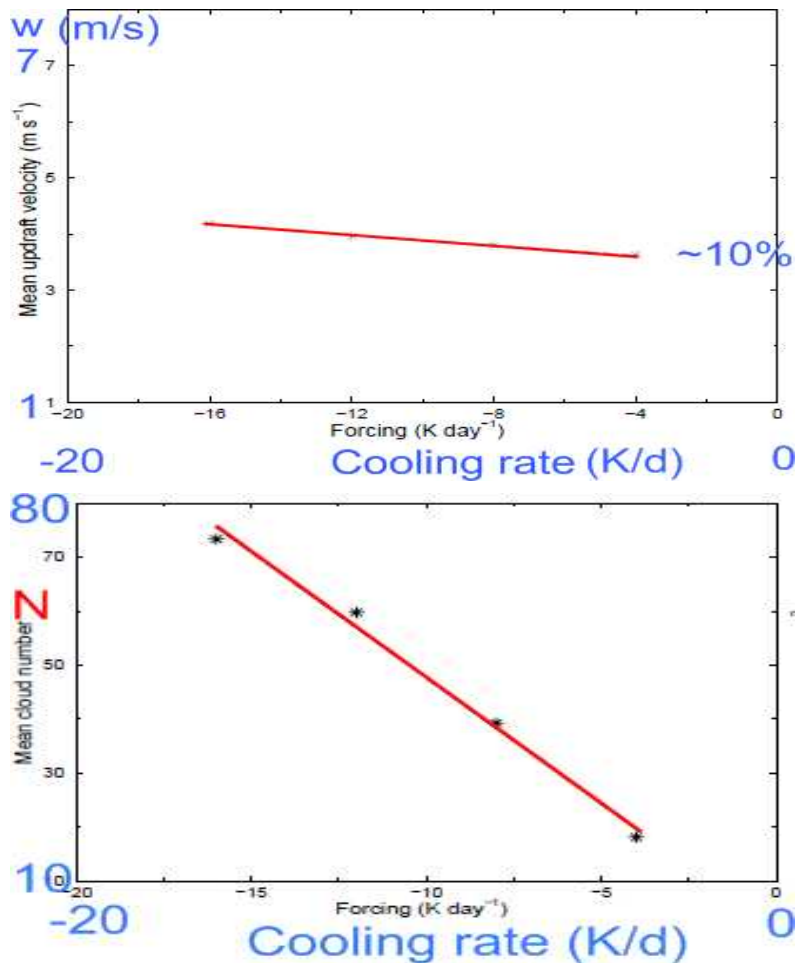
$$K_i \sim M_i^2$$

(Recall  $K_i \sim \sigma_i w_i^2$  and  $M_i = \rho \sigma_i w_i$ )

- For a bulk system, the time dependence is a damped oscillator that approaches equilibrium after a few  $\tau_D$



# But this is wrong!



Increased forcing linearly increases the mass flux,  $\rho\sigma w$

- achieved by increasing cloud number  $\langle N \rangle$
- not the in-cloud velocities
- nor the sizes of clouds

Scalings and CRM data of Emanuel and Bister 1996; Robe and Emanuel 1996; Grant and Brown 1999; Cohen 2001; Parodi and Emanuel 2009

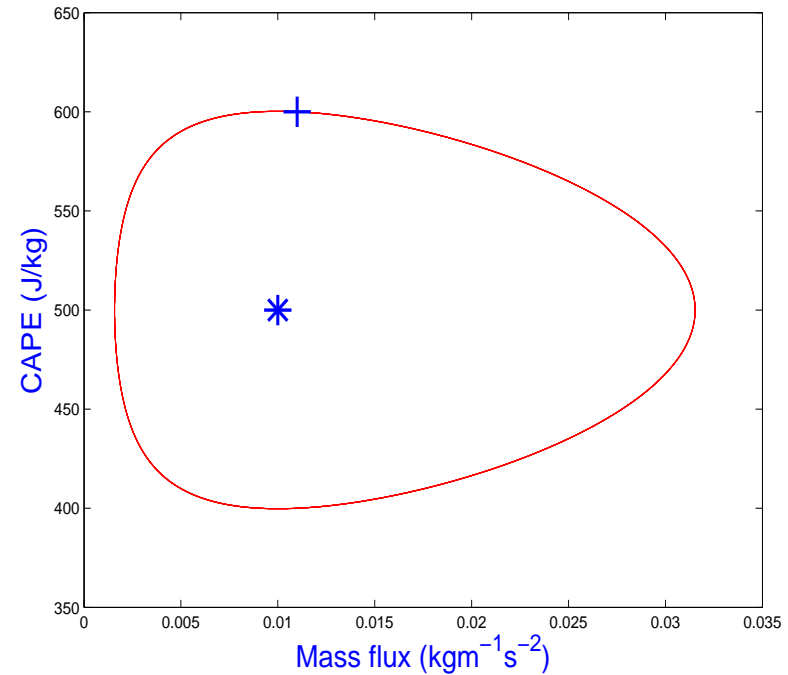
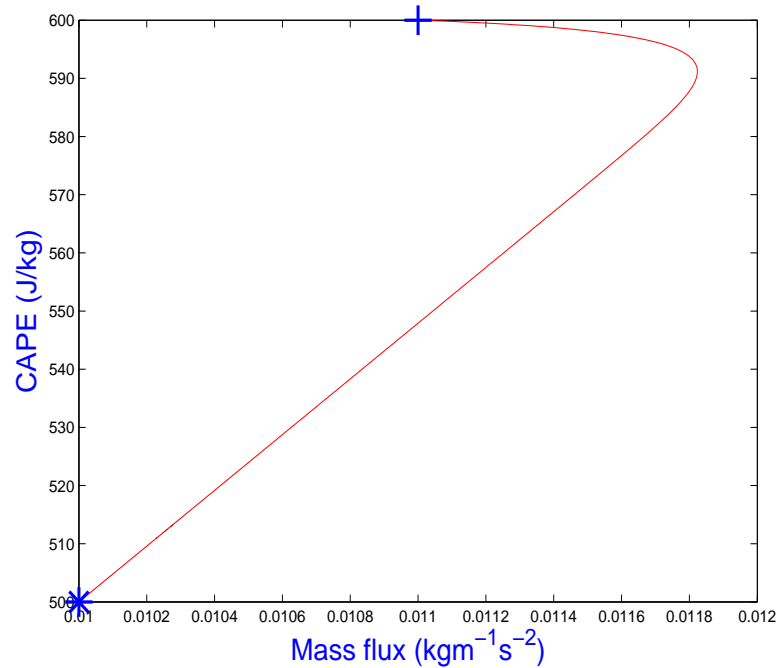
# Closing the system



- To respect the above results, should choose  $K_i \sim M_i$
- Let's consider  $K_i \sim M_i^p$
- Results for a bulk system (one cloud type only)...

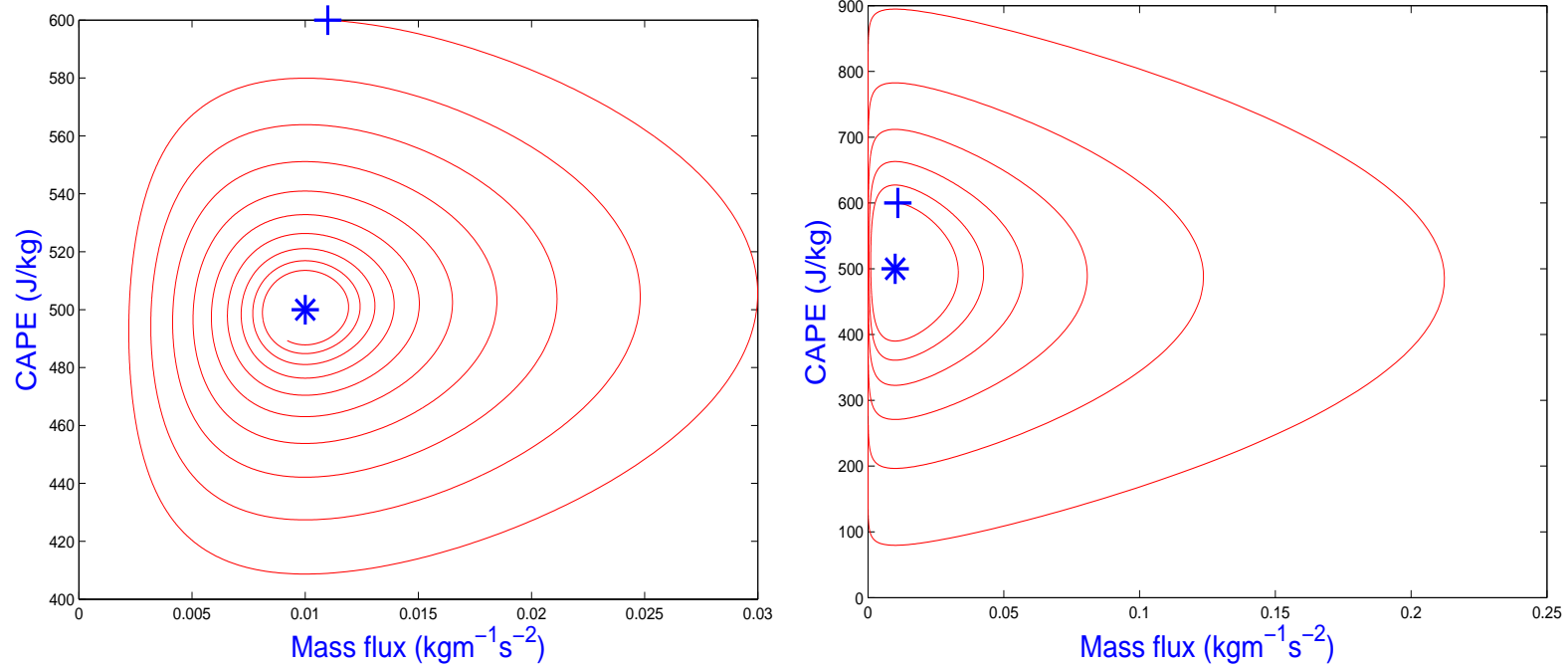


# Illustrative results



$p = 2$  (left) and  $p = 1$  (right) in  $M - B$  space

# Illustrative results



$p = 1.01$  (left) and  $p = 0.99$  (right)

- The CRM data supports  $p \approx 1$  but  $> 1$
- Equilibrium is reached but more slowly as  $p \rightarrow 1$  from above



# Population Dynamics



- Truncate the prognostic system at  $dM_i/dt$  (neglect  $(dM_i/dt)^2$  and  $d^2M_i/dt^2$ ) and write  $M_i = N_i m_i$  to get

$$(p - 1)B_i \frac{dN_i}{dt} = F_i N_i - \sum_j \gamma_{ij} m_j N_i N_j$$

- For  $p > 1$ , a Lotka-Volterra (LV) system of biological populations competing for resource
- i.e., of cloud types competing to remove the instability
- extensively** studied by mathematical ecologists



# Simple Example



- Consider two cloud types, shallow cumulus and deep cumulus
- Described by different parameterization schemes in GCM
- Which one (or both?) to call typically based on ad hoc criteria
- Transitions between them are not well described or understood



# Simple Example



- The LV system has a globally-stable equilibrium state with co-existing shallow and deep clouds if

$$\frac{\gamma_{11}}{\gamma_{12}} < \frac{F_1}{B_2} \quad \text{and} \quad \frac{\gamma_{12}}{\gamma_{22}} < \frac{F_2}{B_1}$$

- Otherwise one type will be driven to extinction
- So in equilibrium-based parameterization we should be using these criteria...





**What is a useful framework  
beyond bulk models of large  $N$ ,  
non-interacting, scale-separated,  
equilibrium systems?**



# Prognostic system for finite $N$ ?



- The prognostic systems above assumed infinite  $N_i$
- Necessary for  $M_i$  to be continuous and  $dM_i/dt$  well defined
- How to generalize to finite  $N$ ?



# Methodology



- Construct an individual-based model with a difference equation for  $P(\{N_i\}, t)$  that evolves according to transition probabilities for births, deaths, competitive exclusion etc
- Choose the probabilities such that in the limit of large system size, we recover the deterministic ode's from before
- Leading correction for a non-infinite system is stochastic and accounts for fluctuations in  $N$
- Explicit demonstration for the biological case has been done (McKane and Newman 2004)  
Straightforward to generalize to a lattice and modulate transition probabilities with  $w_{12}$



# Summary



- The archetypal convective parameterization is based on a bulk model of entraining/detraining plumes
- If grid boxes are not large, fluctuations about statistical equilibrium become important
- If cloud-cloud interactions are important, can account for them if we can say something about  $w_{12}$
- Worthwhile to ask which (or if) equilibrium is reached as this leads to useful constraints
- Proposed framework for a non-equilibrium, spatially-correlated, finite  $N$  model of cumulus
- Could be a useful intermediate system to study, sitting between CRM/observations and parameterization?

